

Hamiltonian formulation of teleparallel gravity using the language of differential forms

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Outline

- 1 3 + 1 decomposition of teleparallel geometry
- 2 New General Relativity action and Lagrangian
- 3 Calculating the New General Relativity Hamiltonian
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- Introduce time translation vector field:
 - Curve through every point $\mathbf{x} \in M$:

$$\begin{aligned} \gamma_{\mathbf{x}} : \mathbb{R} &\rightarrow M \\ t &\mapsto i(\mathbf{t}(\mathbf{x}) + t, \mathbf{s}(\mathbf{x})). \end{aligned}$$

- Vector field at \mathbf{x} :

$$(\partial_t)_{\mathbf{x}} = \left. \frac{d}{dt} \gamma_{\mathbf{x}}(t) \right|_{t=0}.$$

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- **NB! $\partial_{\mathbf{t}}$ is *not* a unit normal vector field - no metric used!**

Split of vector fields and differential forms

- Split of a vector field $\mathbf{X} \in \Gamma(TM)$:

- Temporal part:

$$\check{\mathbf{X}} = \hat{\mathbf{X}}\partial_t = (\mathbf{X} \lrcorner \mathbf{d}t)\partial_t.$$

- Spatial part:

$$\vec{\mathbf{X}} = \mathbf{X} - \check{\mathbf{X}}.$$

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- Split of a differential k -form τ :

- Temporal part:

$$\check{\tau} = \mathbf{d}t \wedge \hat{\tau} = \mathbf{d}t \wedge (\partial_t \lrcorner \tau).$$

- Spatial part:

$$\vec{\tau} = \tau - \check{\tau} = \partial_t \lrcorner (\mathbf{d}t \wedge \tau).$$

Pullback to spatial geometry on Σ

- Definition of time slice Σ_t for $t \in \mathbb{R}$:

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- Time dependent spatial geometry on Σ :

- Pullback of (spatial) vector fields $\mathbf{X} \in \Gamma(T^s M)$:

$$\mathbf{X}(t) = \mathbf{s}_* \circ \mathbf{X} \circ i_t \in \Gamma(T\Sigma).$$

- Pullback of (spatial) differential forms $\tau \in \Omega^{ks}(M)$:

$$\tau(t) = i_t^* \tau \in \Omega^k(\Sigma).$$

Tetrads and teleparallel geometry

- Define the tetrad (coframe) and its dual (frame):

$$\theta^a \in \Omega^1(M), \quad \mathbf{e}_a \in \Gamma(TM), \quad \mathbf{e}_a \lrcorner \theta^b = \delta_a^b.$$

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$$\eta_{ab} = \text{diag}(-1, 1, 1, 1).$$

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- Here: use Weitzenböck gauge $\omega^a_b \equiv 0$.

Space-time split of tetrad geometry

- Decomposition and pullback of the tetrad:

$$\theta^a = \hat{\theta}^a \mathbf{d}\mathbf{t} + \vec{\theta}^a, \quad \hat{\theta}(t) = i_t^* \hat{\theta}^a, \quad \vec{\theta}(t) = i_t^* \vec{\theta}^a.$$

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- Define the basis complement:

$$\xi^a = -\frac{1}{6} \eta^{ae} \epsilon_{abcd} * (\vec{\theta}^b \wedge \vec{\theta}^c \wedge \vec{\theta}^d).$$

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- Properties of the basis complement:

- Orthonormality:

$$\eta_{ab} \xi^a \vec{\theta}^b = 0, \quad \eta_{ab} \xi^a \xi^b = -1.$$

- Completeness:

$$\eta^{ab} = -\xi^a \xi^b + q^{-1}(\vec{\theta}^a, \vec{\theta}^b).$$

- Basis decomposition of the temporal tetrad part:

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Lapse, shift and unit normal vector field

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- Relation between unit normal vector field and basis complement:

$$\xi^a(t) = i_t^* (\mathbf{N} - \theta^a).$$

Irreducible decomposition

- For a time-dependent one-form $\tau^a : \mathbb{R} \rightarrow \Omega^1(\Sigma)$ define:

$$\overset{\circ}{\tau}^a = -\xi^a \xi_b \tau^b,$$

$$\overset{\ominus}{\tau}^a = \frac{1}{2} \left[q^{-1}(\vec{\theta}^a, \vec{\theta}_b) \tau^b - q^{-1}(\vec{\theta}^a, \tau^b) \vec{\theta}_b \right],$$

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- Components are orthogonal and complete:

$$\tau^a \wedge * \sigma_a = \overset{\circ}{\tau}^a \wedge * \overset{\circ}{\sigma}_a + \overset{\ominus}{\tau}^a \wedge * \overset{\ominus}{\sigma}_a + \overset{\oplus}{\tau}^a \wedge * \overset{\oplus}{\sigma}_a + \overset{\otimes}{\tau}^a \wedge * \overset{\otimes}{\sigma}_a.$$

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- Irreducible decomposition of the torsion:

$$\mathcal{V}^a = \frac{1}{3}\theta^a \wedge (\mathbf{e}_b \lrcorner \mathbf{T}^b),$$

$$\mathcal{A}^a = \frac{1}{3}\mathbf{e}^a \lrcorner (\theta_b \wedge \mathbf{T}^b),$$

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Teleparallel action

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- Action with three free constants C_T, C_V, C_A :

$$S[\theta^a] = \int_M \mathcal{L} = \int_M (C_T \mathcal{T}^a \wedge \star \mathcal{T}_a + C_V \mathcal{V}^a \wedge \star \mathcal{V}_a + C_A \mathcal{A}^a \wedge \star \mathcal{A}_a).$$

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- Equivalent expression for the Lagrangian:

$$\begin{aligned} \mathcal{L} &= \frac{2C_T + C_V}{3} \mathbf{T}^a \wedge \star \mathbf{T}_a \\ &+ \frac{C_T - C_V}{3} (\mathbf{T}^a \wedge \theta_b) \wedge \star (\mathbf{T}^b \wedge \theta_a) + \frac{C_A - C_T}{3} (\mathbf{T}^a \wedge \theta_a) \wedge \star (\mathbf{T}^b \wedge \theta_b). \end{aligned}$$

- Decomposition of the Lagrangian:

$$\mathcal{L} = \check{\mathcal{L}} = \mathbf{d}t \wedge \hat{\mathcal{L}} = \mathbf{d}t \wedge (\partial_t \lrcorner \mathcal{L}).$$

Space-time split of the Lagrangian

- Decomposition of the Lagrangian:

$$\mathcal{L} = \check{\mathcal{L}} = \mathbf{d}\mathbf{t} \wedge \hat{\mathcal{L}} = \mathbf{d}\mathbf{t} \wedge (\partial_{\mathbf{t}} - \mathcal{L}).$$

- Pullback to Σ :

$$\begin{aligned} \hat{\mathcal{L}} = & \frac{2C_T + C_V}{3} \left\{ -\frac{1}{\alpha} \left[\dot{\vec{\theta}}^a - \mathbf{d}(\alpha\xi^a) - \boldsymbol{\xi}_\beta \vec{\theta}^a \right] \wedge * \left[\dot{\vec{\theta}}_a - \mathbf{d}(\alpha\xi_a) - \boldsymbol{\xi}_\beta \vec{\theta}_a \right] + \alpha \mathbf{d}\vec{\theta}^a \wedge * \mathbf{d}\vec{\theta}_a \right\} \\ & + \frac{C_T - C_V}{3} \left[-\frac{1}{\alpha} \left(\dot{\vec{\theta}}^a \wedge \vec{\theta}_b + E^a_b \right) \wedge * \left(\dot{\vec{\theta}}^b \wedge \vec{\theta}_a + E^b_a \right) + \alpha (\mathbf{d}\vec{\theta}^a \wedge \vec{\theta}_b) \wedge * (\mathbf{d}\vec{\theta}^b \wedge \vec{\theta}_a) \right] \\ & + \frac{C_A - C_T}{3} \left[-\frac{1}{\alpha} \left(\dot{\vec{\theta}}^a \wedge \vec{\theta}_a + E^a_a \right) \wedge * \left(\dot{\vec{\theta}}^b \wedge \vec{\theta}_b + E^b_b \right) + \alpha (\mathbf{d}\vec{\theta}^a \wedge \vec{\theta}_a) \wedge * (\mathbf{d}\vec{\theta}^b \wedge \vec{\theta}_b) \right] \end{aligned}$$

where

$$E^b_a = -\mathbf{d}(\alpha\xi^b) \wedge \vec{\theta}_a + \alpha\xi_a \mathbf{d}\vec{\theta}^b - (\boldsymbol{\xi}_\beta \vec{\theta}^b) \wedge \vec{\theta}_a.$$

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$$\overset{\odot}{\pi}_a = -\frac{2}{3}(2C_T + C_V)\overset{\odot}{v}_a,$$

$$\overset{\ominus}{\pi}_a = -\frac{2}{3}(2C_A + C_T)\overset{\ominus}{v}_a,$$

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- Solve for velocities and use momenta instead.

General structure of the Hamiltonian

- Bare Hamiltonian $\hat{\mathcal{H}}_0 = v^a \wedge p_a - \hat{\mathcal{L}}$ takes the form

$$\begin{aligned}\hat{\mathcal{H}}_0 &= \hat{\mathcal{H}}_0[\overset{\circ}{p}] + \hat{\mathcal{H}}_0[\overset{\ominus}{p}] + \hat{\mathcal{H}}_0[\overset{\oplus}{p}] + \hat{\mathcal{H}}_0[\overset{\otimes}{p}] \\ &+ \frac{C_A - C_T}{3} \alpha \left[\xi_a \xi_b d\vec{\theta}^a \wedge *d\vec{\theta}^b - d\vec{\theta}^a \wedge \theta_a \wedge *(d\vec{\theta}^b \wedge \theta_b) \right] \\ &+ \frac{C_T - C_V}{3} \alpha (\vec{\theta}_a^\# - d\vec{\theta}^a) \wedge *(\vec{\theta}_b^\# - d\vec{\theta}^b) - C_T \alpha d\vec{\theta}^a \wedge *d\vec{\theta}_a \\ &- (\alpha \xi^a + \beta - \vec{\theta}^a) dp_a - d\vec{\theta}^a \wedge (\beta - p_a) + d \left[(\alpha \xi^a + \beta - \vec{\theta}^a) p_a \right].\end{aligned}$$

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$$\begin{aligned}\hat{\mathcal{H}}_0 &= \hat{\mathcal{H}}_0[\overset{\circ}{\rho}] + \hat{\mathcal{H}}_0[\overset{\ominus}{\rho}] + \hat{\mathcal{H}}_0[\overset{\oplus}{\rho}] + \hat{\mathcal{H}}_0[\overset{\otimes}{\rho}] \\ &+ \frac{C_A - C_T}{3} \alpha \left[\xi_a \xi_b d\vec{\theta}^a \wedge *d\vec{\theta}^b - d\vec{\theta}^a \wedge \theta_a \wedge *(d\vec{\theta}^b \wedge \theta_b) \right] \\ &+ \frac{C_T - C_V}{3} \alpha (\vec{\theta}_a^\# - d\vec{\theta}^a) \wedge *(\vec{\theta}_b^\# - d\vec{\theta}^b) - C_T \alpha d\vec{\theta}^a \wedge *d\vec{\theta}_a \\ &- (\alpha \xi^a + \beta - \vec{\theta}^a) dp_a - d\vec{\theta}^a \wedge (\beta - p_a) + d[(\alpha \xi^a + \beta - \vec{\theta}^a) p_a].\end{aligned}$$

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 - Total derivative (does not contribute).

Terms quadratic in the momenta

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Possible additional primary constraints

- Vector constraint: if $2C_T + C_V = 0$, then

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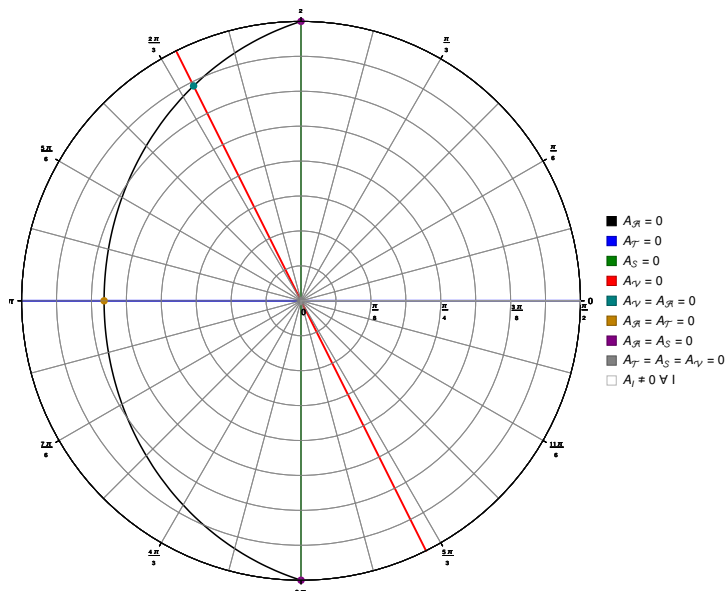
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Visualization of constraints




Outline

- 1 3 + 1 decomposition of teleparallel geometry
- 2 New General Relativity action and Lagrangian
- 3 Calculating the New General Relativity Hamiltonian
- 4 **Conclusion**

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- Application to New General Relativity
 - Use vector-axial-tensor decomposition in the action.
 - Easy to derive, decompose and invert canonical momenta.
 - Kinematic Hamiltonian obtained from Similar decomposition.
 - Appearance of constraints depend on parameters.

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