

# The geometric foundation of gravity

Manuel Hohmann

Laboratory of Theoretical Physics - Institute of Physics - University of Tartu  
Center of Excellence “The Dark Side of the Universe”



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- 1 Introduction
- 2 Building blocks of differential geometry
- 3 A Nobel Prize for geometry
- 4 Three pathways to general relativity
- 5 Going beyond general relativity
- 6 Conclusion

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  - Solar system, planetary motion, binary systems. . .
  - Galaxies, galactic clusters and superclusters, structure formation. . .
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- The geometry of spacetime determines the motion of bodies (matter).
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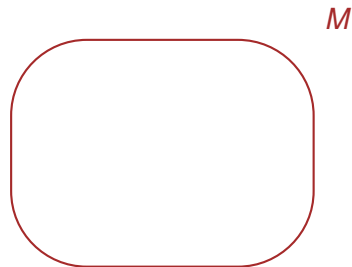
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⇒ Our description of geometry determines how we describe gravity.

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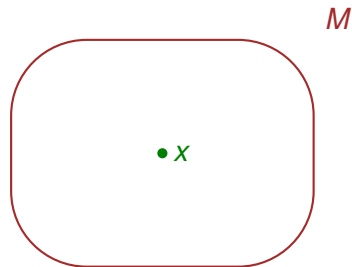
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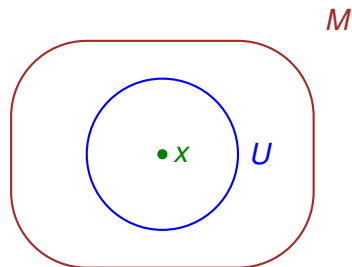
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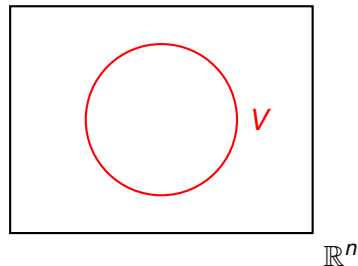
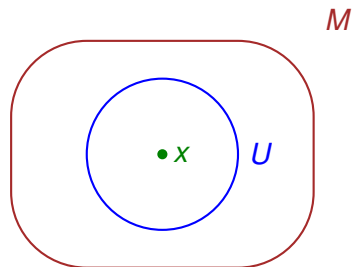
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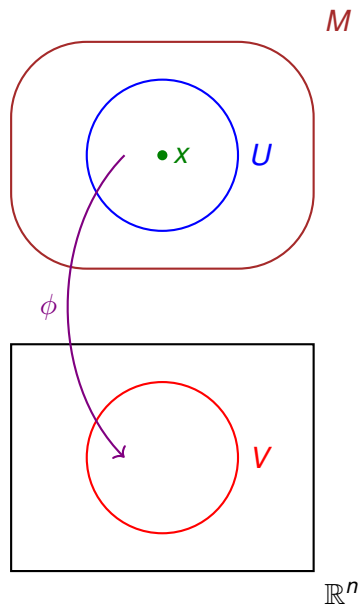
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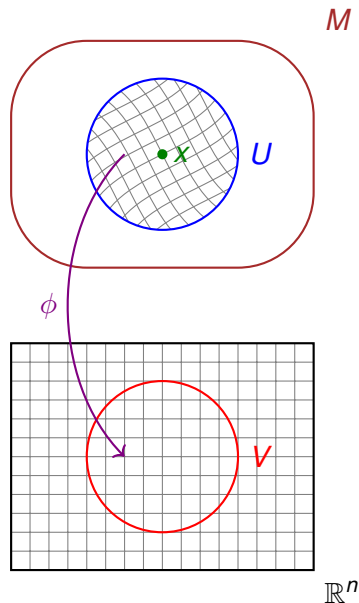
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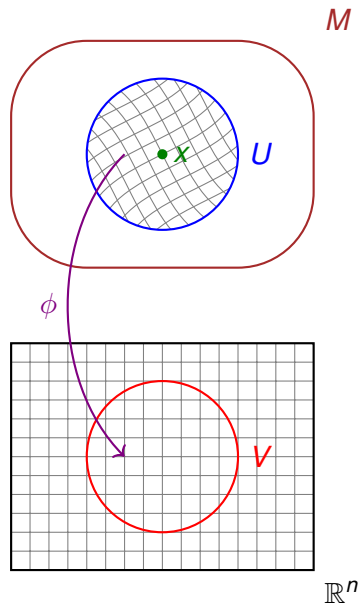


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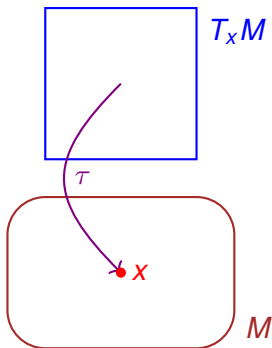
## Notions from differential geometry

- $(U, \phi) \iff$  chart.
- Collection  $\mathcal{A} = \{(U, \phi)\} \iff$  atlas.
- $(M, \mathcal{A}) \iff$  manifold.



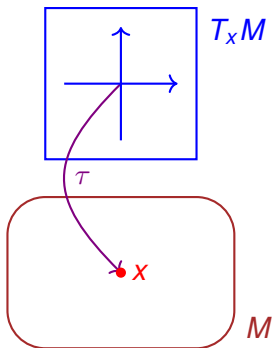
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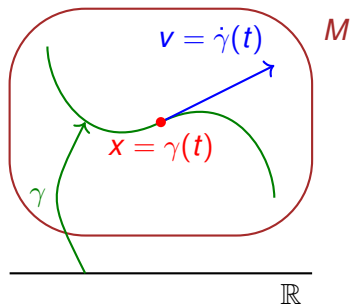
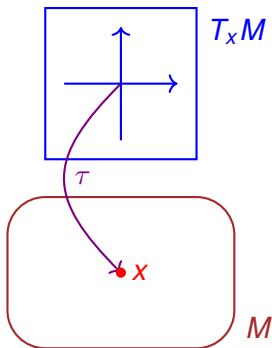
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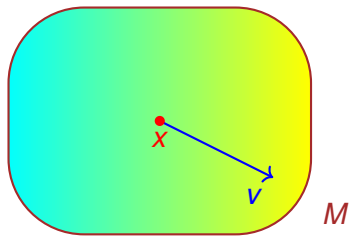
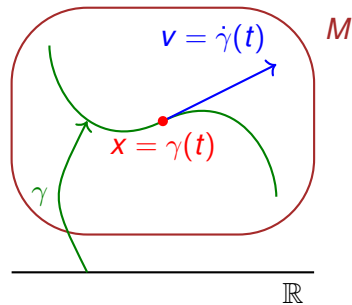
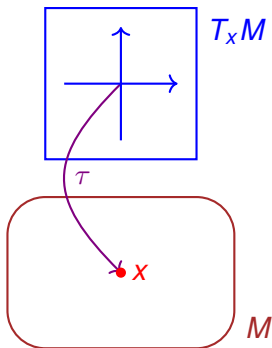
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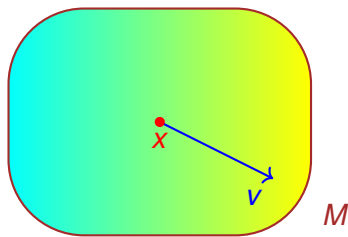
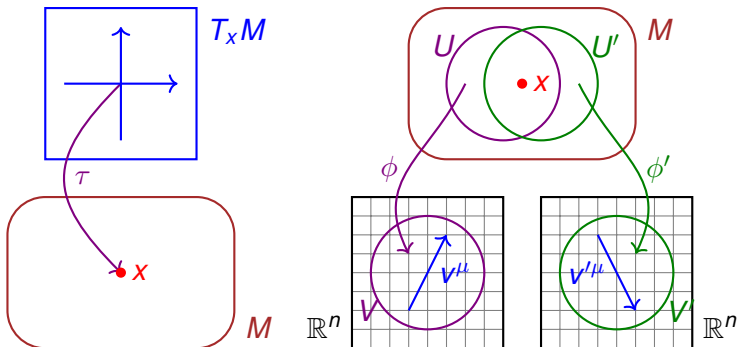
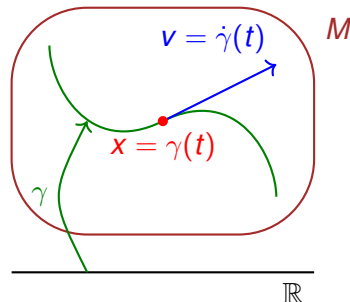
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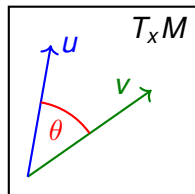
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  3. Components  $v^\mu$  in a chart and their transformation.



# Metric geometry - length, time & causality

- A metric  $g$  is a scalar product on every tangent space  $T_x M$ :

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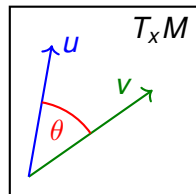
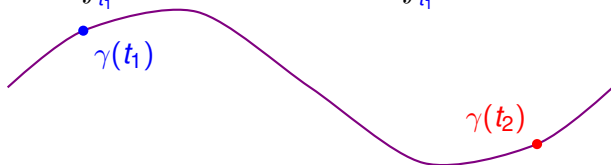
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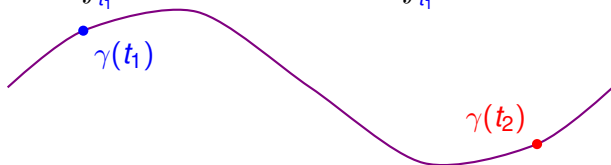
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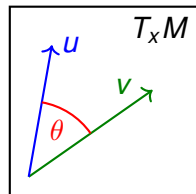
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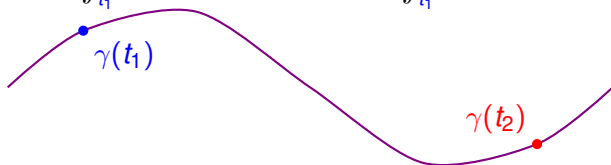
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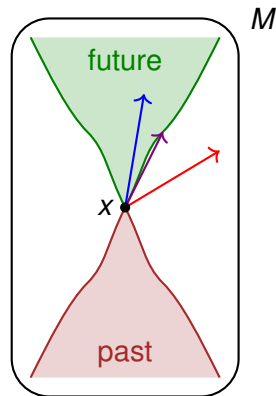
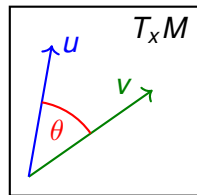
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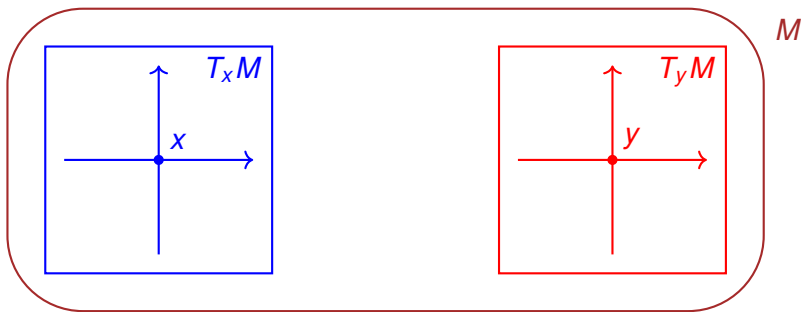
- Length of a trajectory  $\ell \iff$  time measured by moving clock.
- Metric determines causality and propagation of information:

$g(v, v) > 0$	$g(v, v) = 0$	$g(v, v) < 0$
spacelike	lightlike (null)	timelike



# Connections - parallel transport and autoparallel curves

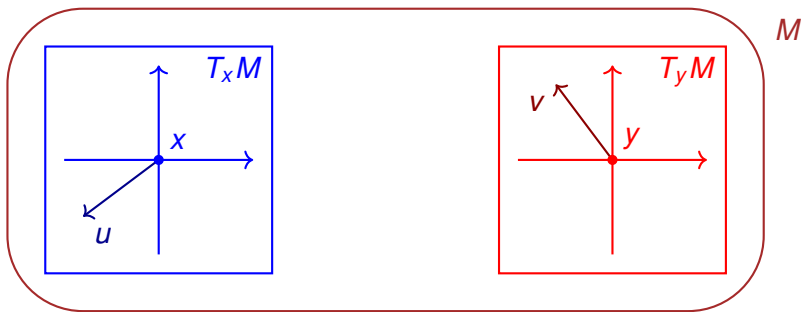
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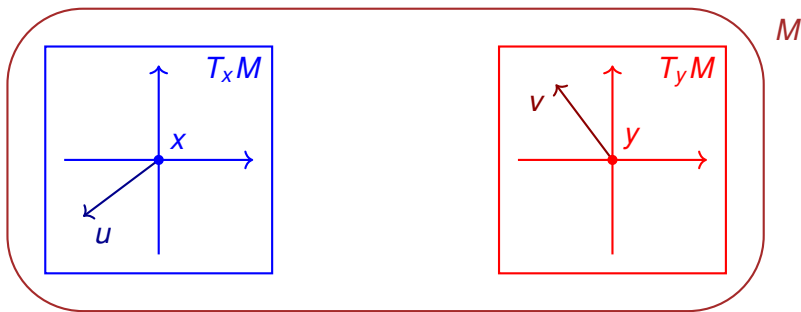
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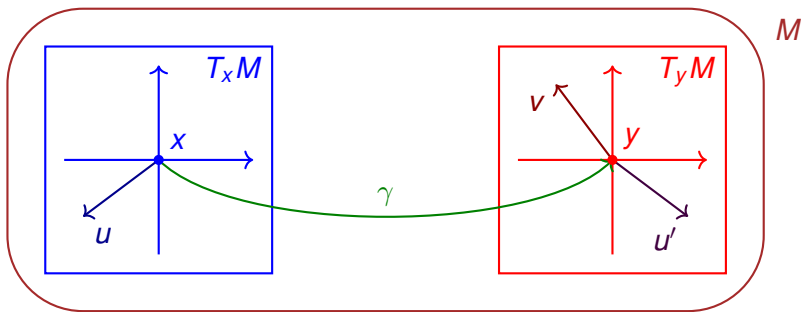
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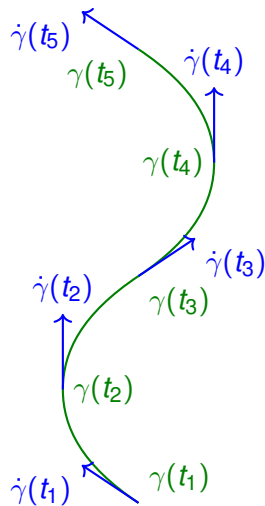
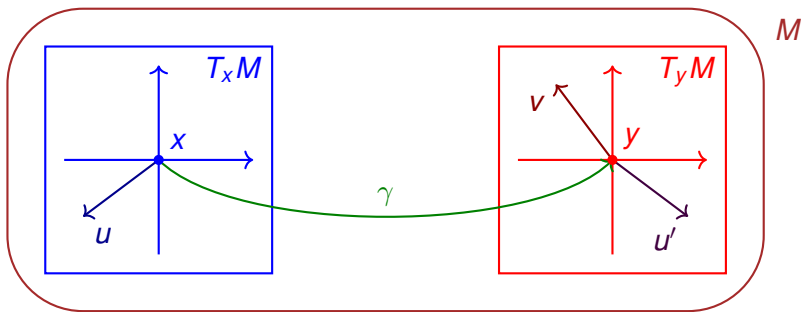
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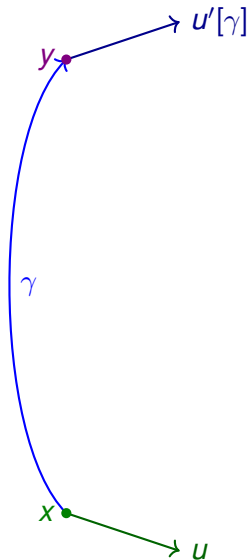
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- Autoparallel curve  $\iff$  parallel transport of tangent vector  $\dot{\gamma}$ .



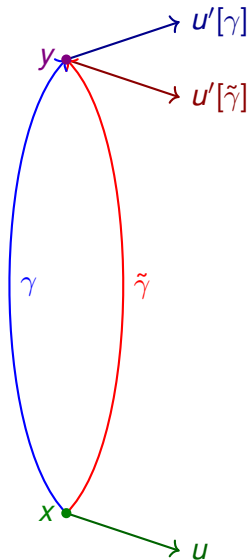
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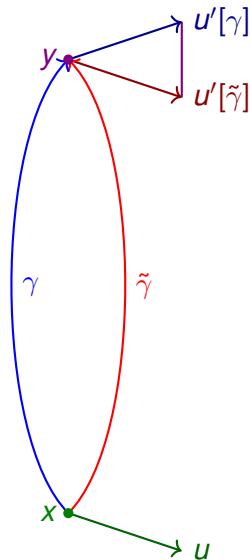
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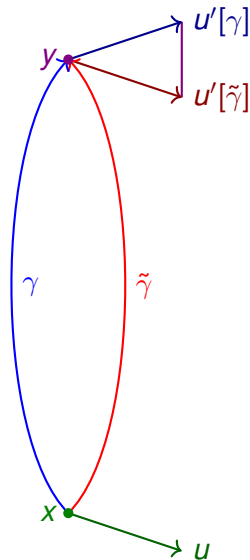
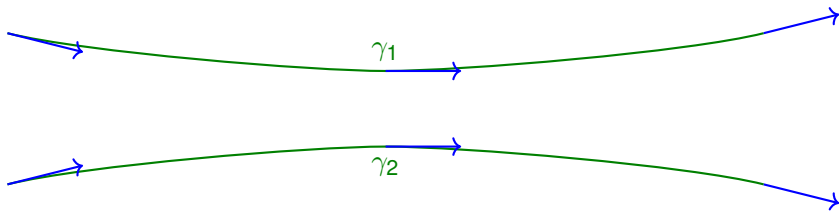
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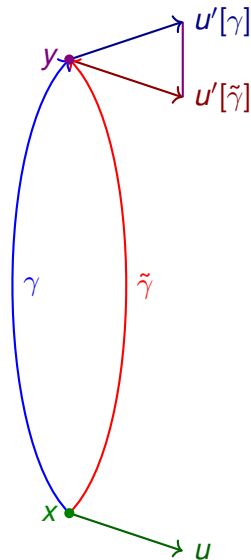
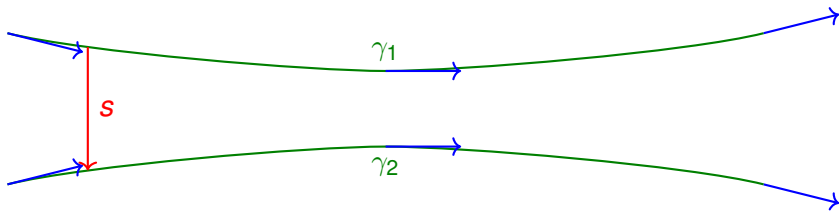
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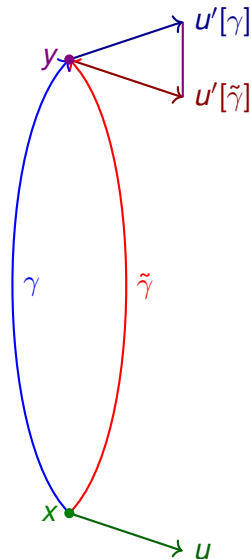
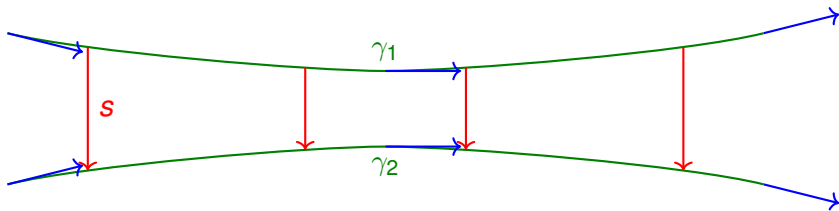
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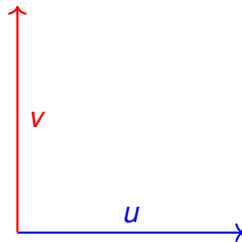
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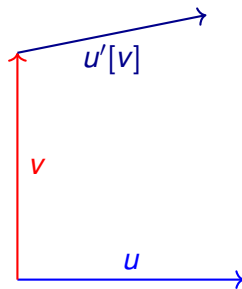
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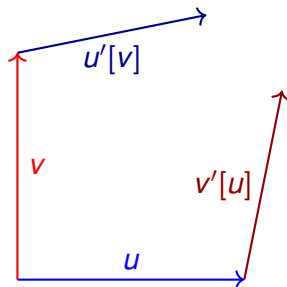
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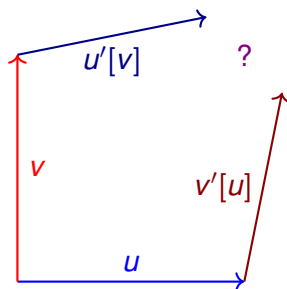
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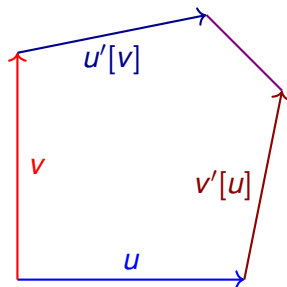
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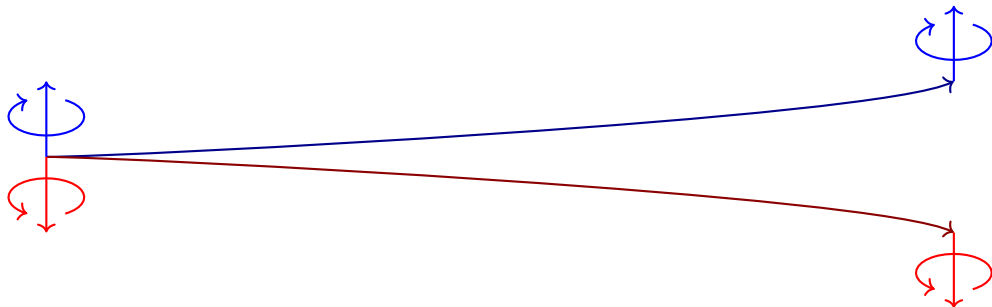
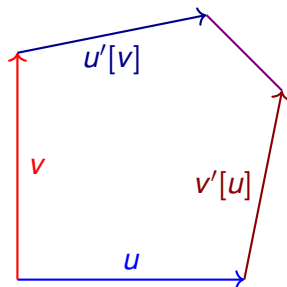
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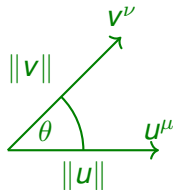
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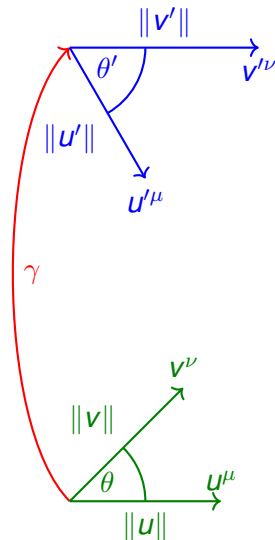
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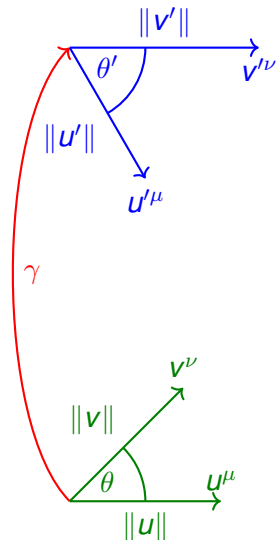
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- Geometric interpretation of  $Q \neq 0$ ?
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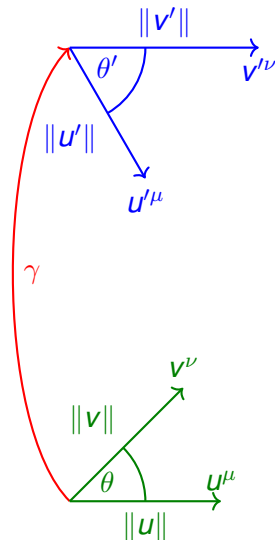
$$g_{\mu\nu} u^{\mu} v^{\nu} \neq g'_{\mu\nu} u'^{\mu} v'^{\nu}.$$

- Length of vectors changes along transport:

$$\|u\| \neq \|u'\|, \quad \|v\| \neq \|v'\|.$$

- Angle between vectors changes along transport:

$$\theta \neq \theta'.$$



# Connection decomposition: Levi-Civita, contortion and disformation

- In presence of a metric, a connection may be uniquely decomposed:

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- 1 Introduction
- 2 Building blocks of differential geometry
- 3 A Nobel Prize for geometry**
- 4 Three pathways to general relativity
- 5 Going beyond general relativity
- 6 Conclusion

# The Nobel Prize in Physics 2020

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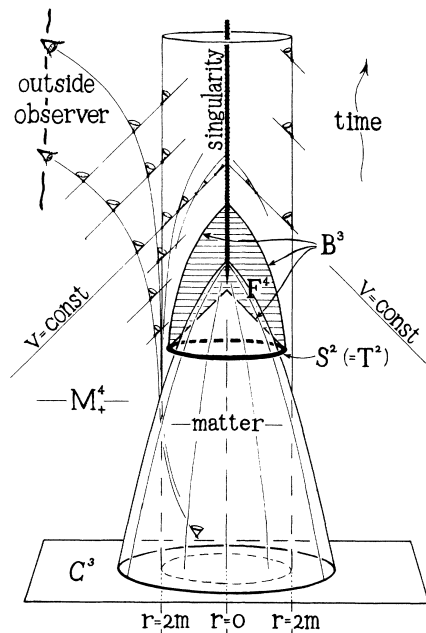
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# Penrose's singularity theorem (1965)

## Singularity theorem:

The formation of a singularity is unavoidable, if:

1. the null energy condition holds,



R. Penrose, Phys. Rev. Lett. **14** (1965) 57.



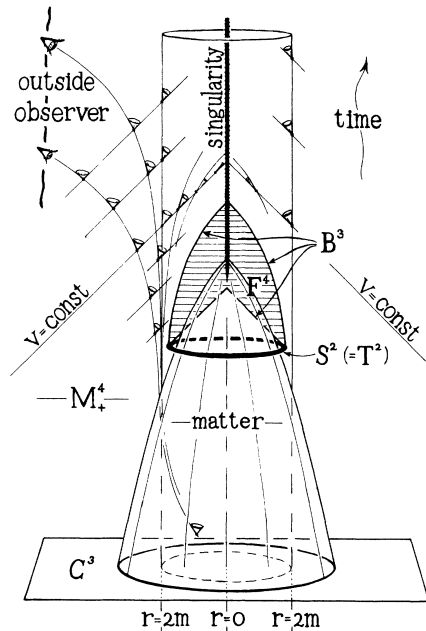
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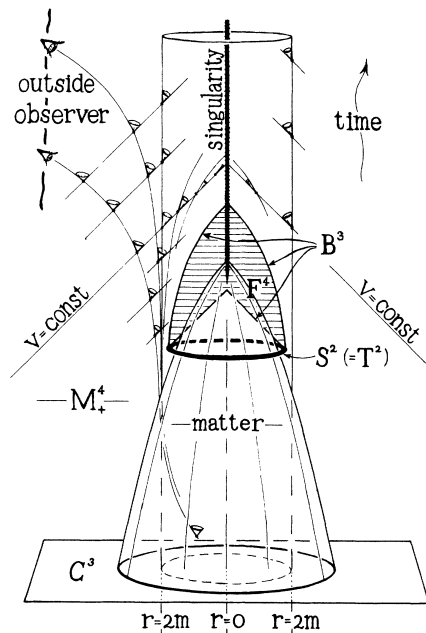
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3. there exist trapped surfaces  $T^2$ ,
4. there exists a Cauchy hypersurface  $C^3$ .

$C^3$  is a Cauchy hypersurface means:

- Every causal trajectory from  $M_+^4$  meets  $C^3$ .
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## • Quantum field theory:

- “Quantum inequalities” hold.
- **Averaged** energy density.

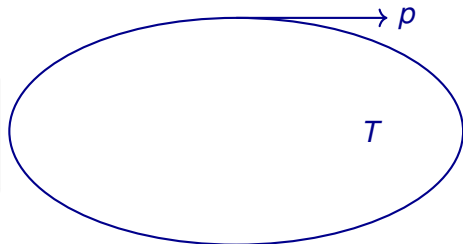


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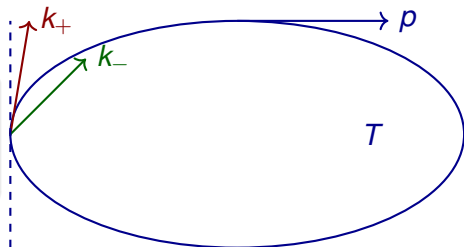


## Geometric description of a trapped surface

- Spacelike surface  $T$ : tangent vectors  $p \parallel T$  are spacelike:  $g(p, p) > 0$ .

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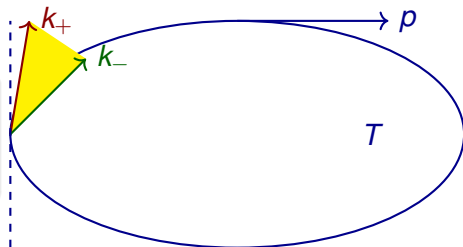
## Geometric description of a trapped surface

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- ⇒ The vectors  $k_{\pm}$  describe the propagation of light; all light from  $T$  moves inwards.

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⚡ In presence of singularities, the fate of free-fall observers is not determinable!

- 1 Introduction
- 2 Building blocks of differential geometry
- 3 A Nobel Prize for geometry
- 4 Three pathways to general relativity**
- 5 Going beyond general relativity
- 6 Conclusion

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- Curvature of the Levi-Civita connection:

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# The “zoo” of modified gravity theories

## Lorentz violation

LV massive grav.

Horava-Lifshitz

ghost condensate

extended HL

Finsler

cuscuton

## Lorentz invariance

higher spin

partially  
massless  
spin 3

## spin 2 gravitation

### massless spin 2

Brans-Dicke

chameleon

$f(R)$

symmetron

Horndeski

galileon

DBI-galileon

multi-galileon

massive  
graviton-  
galileon

### massive spin 2

casca. grav.

DGP

quasi-dilaton

massive grav.

bi-/multi-grav.

# The $f(\dots)$ family of gravity theories

- Action with higher order (curvature, torsion, nonmetricity) terms:
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  - New dynamical effects in cosmology modeling inflation and dark energy.
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- Relation between different extensions?
  - Original Lagrangians differ only by boundary terms:

$$-\mathbb{T} + \overset{\circ}{\nabla}_{\mu} B_T^{\mu} = \overset{\circ}{R} = -\mathbb{Q} + \overset{\circ}{\nabla}_{\mu} B_Q^{\mu}.$$

⇒ Corresponding equivalent theories:

$$f(-\mathbb{T} + \overset{\circ}{\nabla}_{\mu} B_T^{\mu}) = f(\overset{\circ}{R}) = f(-\mathbb{Q} + \overset{\circ}{\nabla}_{\mu} B_Q^{\mu}).$$

- $f(\mathbb{T})$  and  $f(\mathbb{Q})$  Lagrangians lead to essentially different theories.
- Difference cannot be moved into boundary term ⇒ **different field equations.**

# Coupling scalar fields

- Why consider scalar fields  $\Phi$  non-minimally coupled to gravity?
  - Scalar fields are simplest possibility to add another degree of freedom.
  - Discovery of the Higgs boson showed existence of fundamental scalar fields.
  - Scalar fields appear in effective description of other (e.g., string, quantum) theories.
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- Scalar field extensions of different formulations of GR:
  - Scalar-curvature gravity:

$$S_{\text{SCG}} = \frac{1}{16\pi G} \int_M d^4x \sqrt{-\det g} \left[ \mathcal{A}(\Phi) \mathring{R} - \mathcal{B}(\Phi) g^{\mu\nu} \partial_\mu \Phi \partial_\nu \Phi - \mathcal{V}(\Phi) \right].$$

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- Scalar-nonmetricity gravity:

$$S_{\text{SNG}} = \frac{1}{16\pi G} \int_M d^4x \sqrt{-\det g} \left[ -\mathcal{A}(\Phi) \mathbb{Q} - \left( \mathcal{B}(\Phi) \mathring{\nabla}^\mu \Phi - 2\mathcal{C}(\Phi) Q_\nu^{\nu\mu} - 2\mathcal{D}(\Phi) Q^{\mu\nu}{}_\nu \right) \mathring{\nabla}_\mu \Phi - \mathcal{V}(\Phi) \right].$$



# New GR, newer GR and even newer theories

- More general theories constructible from torsion terms:
  - “New general relativity”:

$$\mathcal{L} = c_1 T^{\mu\nu\rho} T_{\mu\nu\rho} + c_2 T^{\mu\nu\rho} T_{\rho\nu\mu} + c_3 T^\mu{}_{\mu\rho} T_\nu{}^{\nu\rho}.$$

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# Non-equivalence of modified gravity theories

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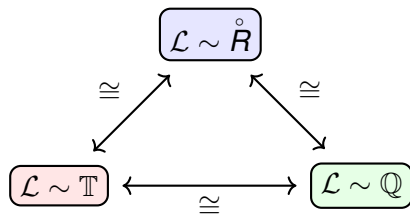
$$\mathcal{L} \sim \overset{\circ}{R}$$

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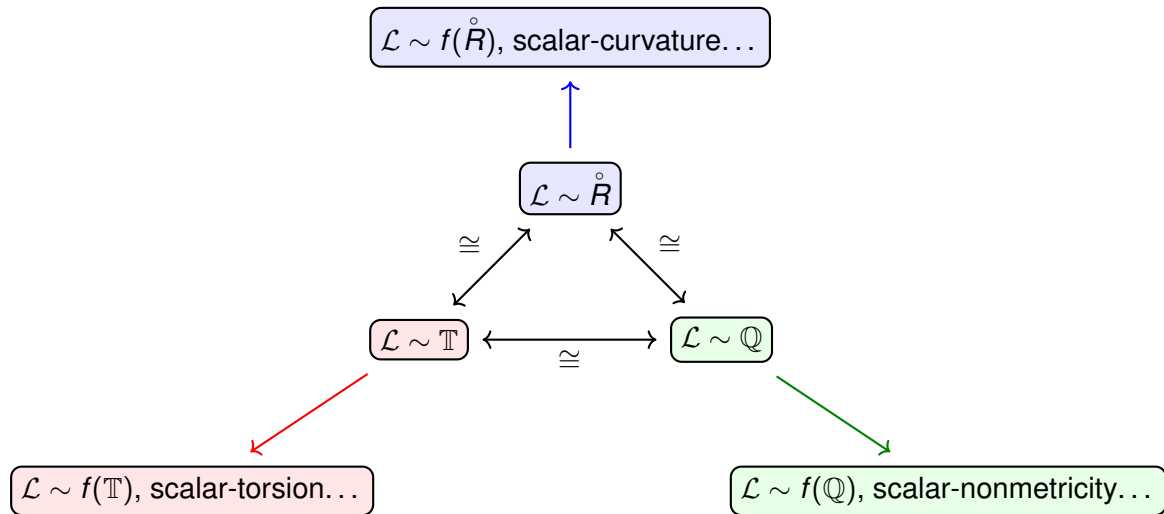
# Non-equivalence of modified gravity theories

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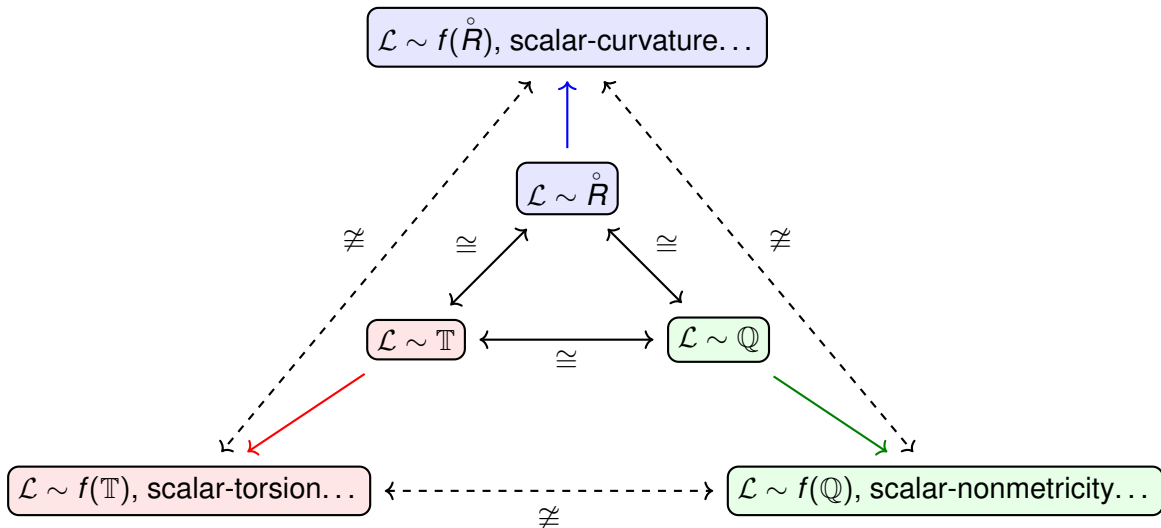
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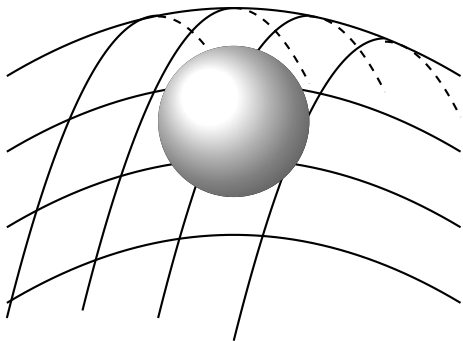
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# A unified picture: Cartan geometry

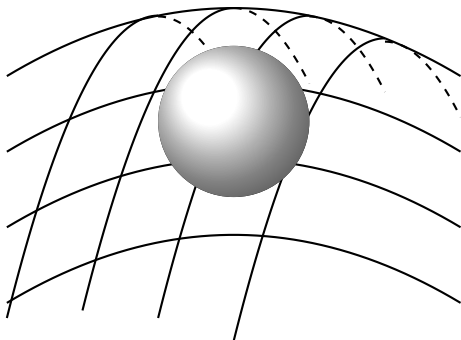
- Cartan geometry: how a hamster sitting in a ball describes geometry.





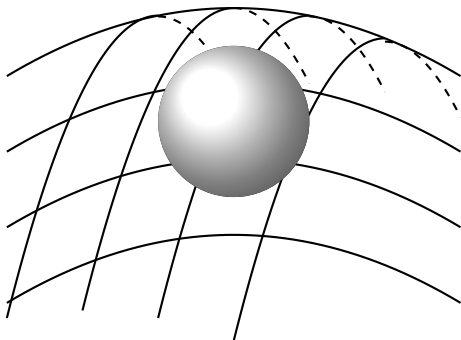
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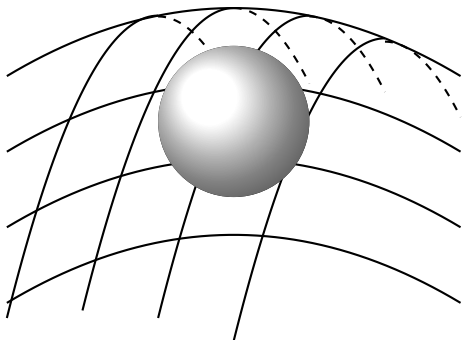
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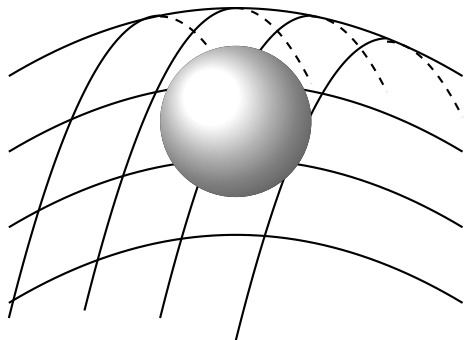
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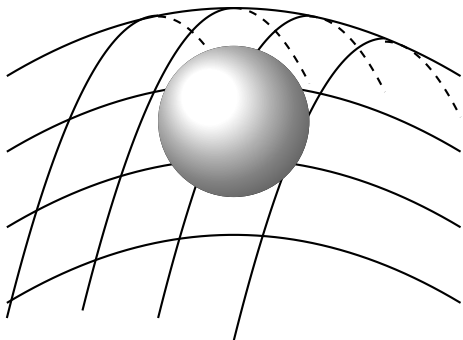


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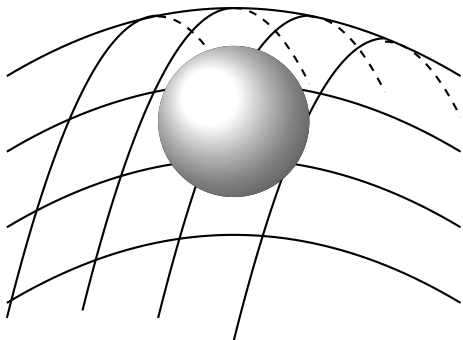
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# The road ahead: from the cosmos to quantum gravity

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  - Gravity as a gauge theory: new similarities with particle physics:
    - Electromagnetism, weak and strong nuclear force modeled by gauge theories.
    - A common description of all forces requires a similar description of gravity.
  - A first order action does not require a boundary term (Gibbons-Hawking-York):
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    - Alternative description of black hole entropy and thermodynamics.
  - A common description using Cartan geometry may pave the path to quantization:
    - Loop Quantum Gravity: canonical quantization of gravity based on Ashtekar variables.
    - Ashtekar variables may be defined naturally in Cartan geometry.

# The road ahead: from the cosmos to quantum gravity

- New geometries provide new insights into well-known problems:
  - How to describe the singularities at the Big Bang and black holes?
  - How to solve the information paradoxes related to black hole horizons?
  - What drives the accelerating expansion of the universe at early and late times?
  - What is the common theory describing all interactions (gravity and particle physics)?
  - How can one construct a consistent theory of quantum gravity?
- What are the advantages of modeling gravity with new geometries?
  - Gravity as a gauge theory: new similarities with particle physics:
    - Electromagnetism, weak and strong nuclear force modeled by gauge theories.
    - A common description of all forces requires a similar description of gravity.
  - A first order action does not require a boundary term (Gibbons-Hawking-York):
    - A boundary term does not enter the field equations, but affects horizons and Casimir effect.
    - Alternative description of black hole entropy and thermodynamics.
  - A common description using Cartan geometry may pave the path to quantization:
    - Loop Quantum Gravity: canonical quantization of gravity based on Ashtekar variables.
    - Ashtekar variables may be defined naturally in Cartan geometry.

⇒ Understanding gravity as geometry is a crucial part of today's physics.