Gravitational wave echoes in multimetric gravity

Manuel Hohmann, Laur Järv, Martti Raidal, Alessandro Strumia, Hardi Veermäe

Laboratory of Theoretical Physics - Institute of Physics - University of Tartu Centre of Excellence "The Dark Side of the Universe"



Centre of Excellence kickoff meeting - 8. March 2016

First observation of gravitational waves



- Black hole merger:
 - Before merger: $36M_{\odot}$ and $29M_{\odot}$ black holes.
 - After merger: $62M_{\odot}$ black hole.
 - \Rightarrow 3*M* $_{\odot}$ radiated into gravitational waves.

- Black hole merger:
 - Before merger: $36M_{\odot}$ and $29M_{\odot}$ black holes.
 - After merger: $62M_{\odot}$ black hole.
 - \Rightarrow 3*M*_{\odot} radiated into gravitational waves.
- \Rightarrow Distance reconstructed from observed amplitude: 410Mpc.
- \Rightarrow Signal travel time to Earth: 1.3Ga.

Idea of wave echoes

• Binary black hole emits different wave components.



Idea of wave echoes

- Binary black hole emits different wave components.
- Different components get deflected differently on their way.



Idea of wave echoes

- Binary black hole emits different wave components.
- Different components get deflected differently on their way.
- Different signal travel times echo.



- Field content:
 - Two metric tensors g¹_{μν}, g²_{μν}.
 Two matter sectors φ¹, φ².

- Field content:
 - Two metric tensors $g^1_{\mu
 u}$, $g^2_{\mu
 u}$.
 - Two matter sectors ϕ^1 , ϕ^2 .
- Structure of the action:

$$S = S_G[g^1, g^2] + S_M[g^1, \phi^1] + S_M[g^2, \phi^2].$$

- Field content:
 - Two metric tensors $g_{\mu\nu}^1, g_{\mu\nu}^2$.
 - Two matter sectors ϕ^1 , ϕ^2 .
- Structure of the action:

$$S = S_G[g^1, g^2] + S_M[g^1, \phi^1] + S_M[g^2, \phi^2]$$
.

- \Rightarrow Both types of matter appear mutually dark.
- \Rightarrow Both metrics may carry gravitational waves.
- \Rightarrow Trajectories of gravitational waves depend on background metrics.

• Metric $g_{\mu\nu}^{l}$ near point mass *M* of type *J*:

$$\begin{split} g_{00}^{I} &= -1 + 2G^{IJ}\frac{M}{r} + \mathcal{O}\left(\frac{M^{2}}{r^{2}}\right) \,, \\ g_{0i}^{I} &= 0 \,, \\ g_{ij}^{I} &= \left(1 + 2G^{IJ}\gamma^{IJ}\frac{M}{r}\right)\delta_{ij} + \mathcal{O}\left(\frac{M^{2}}{r^{2}}\right) \end{split}$$

٠

• Metric $g_{\mu\nu}^{I}$ near point mass *M* of type *J*:

$$\begin{split} g_{00}^{\prime} &= -1 + 2 \boldsymbol{G}^{\prime \prime \prime} \frac{M}{r} + \mathcal{O}\left(\frac{M^2}{r^2}\right) \,, \\ g_{0i}^{\prime} &= 0 \,, \\ g_{ij}^{\prime} &= \left(1 + 2 \boldsymbol{G}^{\prime \prime \prime} \gamma^{\prime \prime \prime} \frac{M}{r}\right) \delta_{ij} + \mathcal{O}\left(\frac{M^2}{r^2}\right) \,. \end{split}$$

- Constants determined by the action:
 - Effective gravitational constants *G^U*.
 - Spatial curvature parameters γ^{IJ} .

• Metric $g_{\mu\nu}^{I}$ near point mass *M* of type *J*:

$$egin{aligned} g_{00}^{\prime} &= -1 + 2G^{\prime J}rac{M}{r} + \mathcal{O}\left(rac{M^2}{r^2}
ight)\,, \ g_{0i}^{\prime} &= 0\,, \ g_{ij}^{\prime} &= \left(1 + 2G^{\prime J}\gamma^{\prime J}rac{M}{r}
ight)\delta_{ij} + \mathcal{O}\left(rac{M^2}{r^2}
ight)\,. \end{aligned}$$

٠

- Constants determined by the action:
 - Effective gravitational constants G^{IJ}.
 - Spatial curvature parameters γ^{IJ} .
- Different spacetime geometries seen by different sectors.

Wave propagation and Shapiro delay

- Assume massless gravitons (lightlike trajectories).
- \Rightarrow Shapiro delay from single deflecting point mass:

$$\delta t = G(1+\gamma) M \ln \left(\frac{(|\vec{x}_{\mathcal{S}}| - \vec{x}_{\mathcal{S}} \cdot \vec{n})(|\vec{x}_{\mathcal{d}}| + \vec{x}_{\mathcal{d}} \cdot \vec{n})}{d^2} \right) \,.$$

- Effective gravitational constant G.
- Spatial curvature parameter γ .
- Source location \vec{x}_s .
- Detector location \vec{x}_d .
- Unit vector \vec{n} from source to detector.
- Shortest distance *d* between wave and deflecting mass.

Wave propagation and Shapiro delay

- Assume massless gravitons (lightlike trajectories).
- \Rightarrow Shapiro delay from single deflecting point mass:

$$\delta t = G(1+\gamma) M \ln \left(\frac{(|\vec{x}_{\mathcal{S}}| - \vec{x}_{\mathcal{S}} \cdot \vec{n})(|\vec{x}_{\mathcal{d}}| + \vec{x}_{\mathcal{d}} \cdot \vec{n})}{d^2} \right) \,.$$

- Effective gravitational constant G.
- Spatial curvature parameter γ .
- Source location \vec{x}_s .
- Detector location \vec{x}_d .
- Unit vector \vec{n} from source to detector.
- Shortest distance *d* between wave and deflecting mass.
- In general: many deflecting masses along the way.
- \Rightarrow Cumulative effect on propagating gravitational wave.

Wave propagation and Shapiro delay

- Assume massless gravitons (lightlike trajectories).
- \Rightarrow Shapiro delay from single deflecting point mass:

$$\delta t = G(1+\gamma) M \ln \left(\frac{(|\vec{x}_{\mathcal{S}}| - \vec{x}_{\mathcal{S}} \cdot \vec{n})(|\vec{x}_{d}| + \vec{x}_{d} \cdot \vec{n})}{d^{2}} \right) \,.$$

- Effective gravitational constant G.
- Spatial curvature parameter γ .
- Source location \vec{x}_s .
- Detector location \vec{x}_d .
- Unit vector \vec{n} from source to detector.
- Shortest distance *d* between wave and deflecting mass.
- In general: many deflecting masses along the way.
- \Rightarrow Cumulative effect on propagating gravitational wave.
- Different parameters G and γ in different sectors.
- Different matter distributions in different sectors.
- \Rightarrow Waves will undergo different Shapiro delays.

Conclusion

- Gravitational wave echo:
 - Assume different gravitational wave components.
 - Gravitational waves get deflected by matter.
 - Different deflection of different components.
 - Different Shapiro delay causes echo.
 - Realization possible via multimetric model.

Conclusion

- Gravitational wave echo:
 - Assume different gravitational wave components.
 - Gravitational waves get deflected by matter.
 - Different deflection of different components.
 - Different Shapiro delay causes echo.
 - Realization possible via multimetric model.
- Calculation of echo parameters:
 - Multimetric action and field equations.
 - Theory parameters G and γ .
 - Matter distribution along wave trajectory.

Conclusion

- Gravitational wave echo:
 - Assume different gravitational wave components.
 - Gravitational waves get deflected by matter.
 - Different deflection of different components.
 - Different Shapiro delay causes echo.
 - Realization possible via multimetric model.
- Calculation of echo parameters:
 - Multimetric action and field equations.
 - Theory parameters G and γ .
 - Matter distribution along wave trajectory.

I have heard the echoes in the dark Dim and distant voices of the past And I've seen so far into the night And lingered in the land of no light. (Uriah Heep, 1991)