

Parameterized post-Newtonian formalism for multimetric gravity

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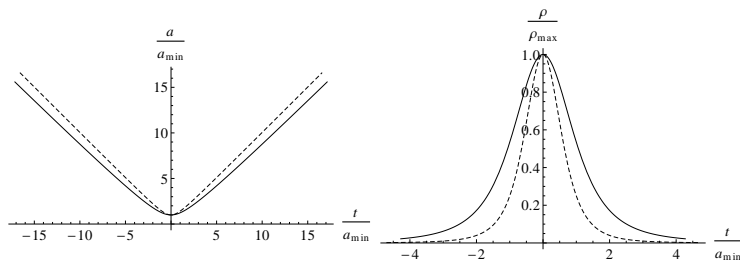
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Motivation

- Λ CDM model: 95% of the universe are dark matter / dark energy.
- Constituents of dark universe are unknown.
- Idea: DM / DE effects from additional *dark* standard model copies.
- Only interaction between standard model copies: repulsive gravity.
- Universe contains equal amounts of matter from all copies.

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- Λ CDM model: 95% of the universe are dark matter / dark energy.
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- Only interaction between standard model copies: repulsive gravity.
- Universe contains equal amounts of matter from all copies.
- Dark galaxies “push” visible matter & light towards visible galaxies.
⇒ Explanation of dark matter!
- Mutual repulsion between galaxies drives accelerating expansion.
⇒ Explanation of dark energy! [MH, M. Wohlfarth '10]



Construction principles of Multimetric gravity

- $N \geq 2$ standard model copies Ψ^I governed by metrics g^I .
- Each standard model copy Ψ^I couples only to its own metric g^I :

$$\Rightarrow S_M[g^I, \Psi^I] = \int d^4x \sqrt{g^I} \mathcal{L}_M[g^I, \Psi^I].$$

- Different sectors couple only gravitationally:

$$\Rightarrow S = S_G[g^1, \dots, g^N] + \sum_{I=1}^N S_M[g^I, \Psi^I].$$

- Field equations obtained from variation with respect to g^I :

$$K_{ab}^I = 8\pi G_N T_{ab}^I$$

- Curvature tensor K_{ab}^I of second derivative order.
- Permutation symmetry of the sectors (g^I, Ψ^I) .
- Vacuum solution given by flat metrics $g^I = \eta$.

- Gravitational field equations:

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Post-Newtonian bookkeeping

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- Slow-moving source matter.
- ⇒ Expand quantities in “velocity” orders $\mathcal{O}(n)$.

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- Matter in the solar system: perfect fluid with
 - density $\rho^I \sim \mathcal{O}(2)$
 - pressure $p^I \sim \mathcal{O}(4)$
 - specific internal energy $\Pi^I \sim \mathcal{O}(2)$
 - velocity $\vec{v}^I \sim \mathcal{O}(1)$

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- Weak gravitational field.

⇒ Expand metric around flat background:

$$g_{ab}^I = \eta_{ab} + h_{ab}^I = \eta_{ab} + h_{ab}^{I(1)} + h_{ab}^{I(2)} + h_{ab}^{I(3)} + h_{ab}^{I(4)}$$

- Each term $h_{ab}^{I(n)}$ is of order $\mathcal{O}(n)$.

- Post-Newtonian metric ansatz:

$$h_{00}^{I(2)} = - \sum_{J=1}^N \alpha^{IJ} \Delta \chi^J,$$

$$h_{\alpha\beta}^{I(2)} = \sum_{J=1}^N \left(2\theta^{IJ} \chi^J_{,\alpha\beta} - (\gamma^{IJ} + \theta^{IJ}) \Delta \chi^J \delta_{\alpha\beta} \right),$$

$$h_{0\alpha}^{I(3)} = \sum_{J=1}^N \left(\sigma_+^{IJ} W_\alpha^{J+} + \sigma_-^{IJ} W_\alpha^{J-} \right),$$

$$h_{00}^{I(4)} = \sum_{J=1}^N \left(\phi_p^{IJ} \Phi_p^J + \phi_{\Pi}^{IJ} \Phi_{\Pi}^J + \sum_{A=1}^2 \omega_A^{IJ} \Omega_A^J \right) + \sum_{J,K=1}^N \sum_{A=1}^7 \psi_A^{IJK} \Psi_A^{JK}.$$

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- Parameters $\alpha^{IJ}, \gamma^{IJ}, \theta^{IJ}, \sigma_{\pm}^{IJ}, \phi_p^{IJ}, \phi_{\Pi}^{IJ}, \omega_1^{IJ}, \omega_2^{IJ}, \psi_1^{IJK}, \dots, \psi_7^{IJK}$.

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- Parameters $\alpha^{IJ}, \gamma^{IJ}, \theta^{IJ}, \sigma_{\pm}^{IJ}, \phi_p^{IJ}, \phi_{\Pi}^{IJ}, \omega_1^{IJ}, \omega_2^{IJ}, \psi_1^{IJK}, \dots, \psi_7^{IJK}$.
- Potentials $\chi^I, W^{I\pm}, \Phi_p^I, \Phi_{\Pi}^I, \Omega_1^I, \Omega_2^I, \Psi_1^{IJ}, \dots, \Psi_7^{IJ}$.

PPN potentials - part 1

- Superpotential:

$$\chi^I = - \int \rho^{II} |\vec{x} - \vec{x}'| d^3 x'.$$

- Vector potentials:

$$W_{\alpha}^{\pm I} = \int \rho^{II} \left(\frac{v_{\alpha}^{II}}{|\vec{x} - \vec{x}'|} \pm \frac{(x_{\alpha} - x'_{\alpha})(x_{\beta} - x'_{\beta}) v_{\beta}^{II}}{|\vec{x} - \vec{x}'|^3} \right) d^3 x'.$$

- Pressure:

$$\Phi_{\rho}^I = \int \frac{\rho^{II}}{|\vec{x} - \vec{x}'|} d^3 x'.$$

- Internal energy:

$$\Phi_{\Pi}^I = \int \frac{\rho^{II} \Pi^{II}}{|\vec{x} - \vec{x}'|} d^3 x'.$$

- Kinetic energy:

$$\Omega_1^I = \int \frac{\rho'^I v'^I{}^2}{|\vec{x} - \vec{x}'|} d^3x', \quad \Omega_2^I = \int \frac{\rho'^I [\vec{v}'^I \cdot (\vec{x} - \vec{x}')]^2}{|\vec{x} - \vec{x}'|^3} d^3x'.$$

- Non-linear potentials:

$$\begin{aligned} \Delta\Delta\Psi_1^{IJ} &= \Delta\chi^I\Delta\Delta\Delta\chi^J, & \Delta\Delta\Psi_2^{IJ} &= \chi^I_{,\alpha\beta}\Delta\Delta\chi^J_{,\alpha\beta}, \\ \Delta\Delta\Psi_3^{IJ} &= \Delta\chi^I_{,\alpha}\Delta\Delta\chi^J_{,\alpha}, & \Delta\Delta\Psi_4^{IJ} &= \chi^I_{,\alpha\beta\gamma}\Delta\chi^J_{,\alpha\beta\gamma}, \\ \Delta\Delta\Psi_5^{IJ} &= \Delta\Delta\chi^I\Delta\Delta\chi^J, & \Delta\Delta\Psi_6^{IJ} &= \Delta\chi^I_{,\alpha\beta}\Delta\chi^J_{,\alpha\beta}, \\ & & \Delta\Delta\Psi_7^{IJ} &= \chi^I_{,\alpha\beta\gamma\delta}\chi^J_{,\alpha\beta\gamma\delta}. \end{aligned}$$

Relation to standard PPN parameters

- PPN parameters relevant for visible matter T_{ab}^1 and metric g_{ab}^1 :

$$\begin{aligned}\alpha^{11} &= \alpha, & \sigma_-^{11} &= -\frac{1}{2} - \gamma - \frac{1}{4}\alpha_1 + \frac{1}{2}\alpha_2 - \frac{1}{2}\zeta_1 + \xi, \\ \gamma^{11} &= \gamma, & \sigma_+^{11} &= -1 - \gamma - \frac{1}{4}\alpha_1, & \phi_{\Pi}^{11} &= 2 + 2\zeta_3, \\ \phi_p^{11} &= 6\gamma + 6\zeta_4 + 4\xi, & \omega_1^{11} &= 2 + 2\gamma + \alpha_3 + \zeta_1 - 2\xi, \\ \omega_2^{11} &= 2\xi - \zeta_1, & \psi_2^{111} &= 2\xi, & \psi_6^{111} &= 6\xi - 2\beta, \\ \psi_1^{111} &= \psi_5^{111} = \frac{1}{2} + \frac{3}{2}\gamma - 2\beta + \frac{1}{2}\zeta_2 + \xi, & \psi_4^{111} &= 4\xi, \\ \psi_3^{111} &= 1 + 3\gamma - 6\beta + \zeta_2 + 6\xi, & \theta^{11} &= \psi_7^{111} = 0.\end{aligned}$$

Relation to standard PPN parameters

- PPN parameters relevant for visible matter T_{ab}^1 and metric g_{ab}^1 :

$$\begin{aligned}\alpha^{11} = \alpha &= \mathbf{1}, & \sigma_-^{11} &= -\frac{1}{2} - \gamma - \frac{1}{4}\alpha_1 + \frac{1}{2}\alpha_2 - \frac{1}{2}\zeta_1 + \xi = -\frac{\mathbf{3}}{\mathbf{2}}, \\ \gamma^{11} = \gamma &= \mathbf{1}, & \sigma_+^{11} &= -1 - \gamma - \frac{1}{4}\alpha_1 = -\mathbf{2}, & \phi_{\Pi}^{11} &= 2 + 2\zeta_3 = \mathbf{2}, \\ \phi_p^{11} &= 6\gamma + 6\zeta_4 + 4\xi = \mathbf{6}, & \omega_1^{11} &= 2 + 2\gamma + \alpha_3 + \zeta_1 - 2\xi = \mathbf{4}, \\ \omega_2^{11} &= 2\xi - \zeta_1 = \mathbf{0}, & \psi_2^{111} &= 2\xi = \mathbf{0}, & \psi_6^{111} &= 6\xi - 2\beta = -\mathbf{2}, \\ \psi_1^{111} = \psi_5^{111} &= \frac{1}{2} + \frac{3}{2}\gamma - 2\beta + \frac{1}{2}\zeta_2 + \xi = \mathbf{0}, & \psi_4^{111} &= 4\xi = \mathbf{0}, \\ \psi_3^{111} &= 1 + 3\gamma - 6\beta + \zeta_2 + 6\xi = -\mathbf{2}, & \theta^{11} = \psi_7^{111} &= \mathbf{0}.\end{aligned}$$

- Compare with measured values of standard PPN parameters.
 $\Rightarrow \theta^{11} = 0$ and $\psi_1^{111} = \psi_5^{111}$ due to gauge freedom.
 $\Rightarrow \alpha^{11} = 1$ can be achieved by choice of units.
 \Rightarrow **13 physical parameters accessible to visible matter experiments.**

Example action - part 1

Generic, simple, multimetric gravity action:

$$S_G = \frac{1}{16\pi} \int d^4x \sqrt{g_0} \left[\sum_{l=1}^N \left(c_1 R^l + g^{lij} \left(c_3 \tilde{S}^l{}_i \tilde{S}^l{}_j + c_5 \tilde{S}^l{}_k \tilde{S}^{lk}{}_{ij} + c_7 \tilde{S}^{lk}{}_{il} \tilde{S}^{ll}{}_{jk} \right) + g^{lij} g^{lkl} g^l{}_{mn} \left(c_9 \tilde{S}^{lm}{}_{ik} \tilde{S}^{ln}{}_{jl} + c_{11} \tilde{S}^{lm}{}_{ij} \tilde{S}^{ln}{}_{kl} \right) \right) + \sum_{l,J=1}^N \left(c_2 g^{lJ} R^J{}_{ij} + g^{lij} \left(c_4 S^{lJ}{}_i S^{lJ}{}_j + c_6 S^{lJ}{}_k S^{lJk}{}_{ij} + c_8 S^{lJk}{}_{il} S^{lJl}{}_{jk} \right) + g^{lij} g^{lkl} g^l{}_{mn} \left(c_{10} S^{lJm}{}_{ik} S^{lJn}{}_{jl} + c_{12} S^{lJm}{}_{ij} S^{lJn}{}_{kl} \right) \right) \right].$$

Example action - part 2

- Ricci tensor and Ricci scalar:

$$R^I, \quad g^{Iab} R^J_{ab}.$$

- Connection difference tensors:

$$S^{IJi}_{jk} = \Gamma^{Ii}_{jk} - \Gamma^{Ji}_{jk}, \quad S^{IJ}_j = S^{IJk}_{jk},$$

$$\tilde{S}^{Ji}_{jk} = \frac{1}{N} \sum_{I=1}^N S^{IJi}_{jk}, \quad \tilde{S}^J_j = \tilde{S}^{Jk}_{jk}.$$

- Mixed volume form

$$g_0 = \prod_{I=1}^N (g^I)^{\frac{1}{N}}.$$

- 12 free, constant parameters:

$$c_1, \dots, c_{12}.$$

Repulsive Newtonian limit

- All diagonal elements α^{II} are equal.
- Units can be chosen to rescale $\alpha^{II} = 1$.
- All off-diagonal elements $\alpha^{IJ} = z$ are equal.
- Parameter values:

$$\alpha^{IJ} = \frac{\bar{\alpha}}{N} + \hat{\alpha}\delta^{IJ}$$

with

$$\bar{\alpha} = Nz, \quad \hat{\alpha} = 1 - z.$$

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- Newtonian metric perturbation:

$$\begin{aligned} h_{00}^{I(2)} &= -\Delta \chi^I - z \sum_{J \neq I} \Delta \chi^J \\ &= 2U^I + 2z \sum_{J \neq I} U^J. \end{aligned}$$

⇒ Repulsive Newtonian limit for $z = -1$.

Results

- 6 constants $c_4, c_6, c_8, c_{10}, c_{11}, c_{12}$ remain free parameters.
- 6 constants $c_1, c_2, c_3, c_5, c_7, c_9$ fixed depending on N and z .
- **Multimetric gravity compatible with experiments.**

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- 6 constants $c_1, c_2, c_3, c_5, c_7, c_9$ fixed depending on N and z .
- Multimetric gravity compatible with experiments.
- PPN parameters (+ lengthy expressions for $\psi_1^{IJK}, \dots, \psi_7^{IJK}$):

$$\begin{aligned}\bar{\alpha} &= Nz, & \hat{\alpha} &= 1 - z, \\ \bar{\gamma} &= Nz, & \hat{\gamma} &= 1 - z, \\ \bar{\theta} &= 0, & \hat{\theta} &= 0, \\ \bar{\sigma}_+ &= -2Nz, & \hat{\sigma}_+ &= 2(z - 1), \\ \bar{\sigma}_- &= \frac{N}{2}(1 - 4z), & \hat{\sigma}_- &= 2(z - 1), \\ \bar{\omega}_1 &= N(5z - 1), & \hat{\omega}_1 &= 5(1 - z), \\ \bar{\omega}_2 &= N(1 - z), & \hat{\omega}_2 &= z - 1, \\ \bar{\phi}_\rho &= 2N(4z - 1), & \hat{\phi}_\rho &= 8(1 - z), \\ \bar{\phi}_\Pi &= 2Nz, & \hat{\phi}_\Pi &= 2(1 - z).\end{aligned}$$

- Multimetric PPN formalism:
 - Assume perfect fluid matter.
 - Determine post-Newtonian metric.
 - PPN parameters $\alpha^{IJ}, \gamma^{IJ}, \theta^{IJ}, \sigma_{\pm}^{IJ}, \phi_{\rho}^{IJ}, \phi_{\Pi}^{IJ}, \omega_1^{IJ}, \omega_2^{IJ}, \psi_1^{IJK}, \dots, \psi_7^{IJK}$.
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 - 13 parameters accessible to visible matter experiments.
 - Extension of standard PPN formalism to $N \geq 2$ metrics.
- Application to multimetric repulsive gravity:
 - Simple model dependent on 12 constant parameters.
 - 6 parameters fixed by experimental consistency.
 - Experimentally consistent model with 6 free parameters.
 - Repulsive gravity allows for accelerating expansion.

- Experimental significance of new visible PPN parameters:
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- Extend other formalisms to multimetric gravity:
 - Strong fields and pulsars: parameterized post-Keplerian formalism?
 - Gravitational waves: parameterized post-Einsteinian formalism?